# Weakened Mannheim Curves in Pseudo-Galilean 3-Space 

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#### Abstract

In this study, Frenet-Mannheim curves and Weakened Mannheim curves are investigated in pseudo-Galilean 3-space. Some characterizations for this curves are obtained.


Keywords : Mannheim cuves; Frenet-Mannheim curves; Weakened-Mannheim curves; pseudo-Galilean 3-space.
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## 1 Introduction

In the study of the fundamental theory and the characterizations of space curves, the corresponding relations between the curves are the very interesting and important problem. The well-known Bertrand curve is characterized as a kind of such corresponding relation between the two curves. For the Bertrand curve $\alpha$, it shares the normal lines with another curve $\beta$, called Bertrand mate or Bertrand partner curve of $\alpha$ [1].

In 1967, Lai investigated the properties of two types of similar curves (the Frenet-Bertrand curves and the Weakened Bertrand curves) under weakened conditions [2].

In recent works, Liu and Wang [1, 3] studied the Mannheim curves in both Euclidean and Minkowski 3-space and they obtained the necessary and sufficient conditions between the curvature and the torsion for a curve to be the Mannheim

[^0]partner curves. Meanwhile, the detailed discussion concerned with the Mannheim curves can be found in literature (see Wang and Liu [3, Liu and Wang [1, Orbay and Kasap (4) and references therein [4]. Karacan and Tuncer investigated the properties of two types of similar curves (the Frenet-Mannheim curves and the Weakened Mannheim curves) under weakened conditions, in [5, 6]. Öztekin investigated Weakened Bertrand curves in [7] under weakened conditions.

In this paper, our main purpose is to extend some results which were given in [2] to Frenet-Mannheim curves and Weakened Mannheim curves in pseudoGalilean 3 -space and we assume that, the angle between tangent vectors $T_{\beta}$ and $T_{\alpha}$ is constant such that $\left\langle T_{\alpha}, T_{\beta}\right\rangle=\cosh \theta \neq 0$.

## 2 Preliminaries

The geometry of the pseudo-Galilean space is similar (but not the same) to the Galilean space. The pseudo-Galilean space $G_{3}^{1}$ is a three-dimensional projective space in which the absolute consists of a real plane $w$ (the absolute plane), a real line $f \subset w$ (the absolute line) and a hyperbolic involution on $f$. Projective transformations which preserve the absolute form of a group $H_{8}$ and are in nonhomogeneous coordinates can be written in the form

$$
\begin{align*}
\bar{x} & =a+b x  \tag{2.1}\\
\bar{y} & =c+d x+r \cosh \theta \cdot y+r \sinh \theta \cdot z \\
\bar{z} & =e+f x+r \sinh \theta \cdot y+r \cos \theta \cdot z
\end{align*}
$$

where $a, b, c, d, e, f, r$ and $\theta$ are real numbers. Particularly, for $b=r=1$, the group (2.1) becomes the group $B_{6} \subset H_{8}$ of isometries (proper motions) of the pseudoGalilean space $G_{3}^{1}$. The motion group remains invariant the absolute figure and defines the other invariants of this geometry. It has the following form

$$
\begin{align*}
\bar{x} & =a+x  \tag{2.2}\\
\bar{y} & =c+d x+\cosh \theta \cdot y+\sinh \theta \cdot z \\
\bar{z} & =e+f x+\sinh \theta \cdot y+\cos \theta \cdot z
\end{align*}
$$

According to the motion group in the pseudo-Galilean space, there are nonisotropic vectors $X(x, y, z)$ (for which holds $x \neq 0)$ and four types of isotropic vectors: spacelike $\left(x=0, y^{2}-z^{2}>0\right)$, timelike $\left(x=0, y^{2}-z^{2}<0\right)$ and two types of lightlike vectors $(x=0, y=\mp z)$. The scalar product of two vectors $A=\left(a_{1}, a_{2}, a_{3}\right)$ and $B=\left(b_{1}, b_{2}, b_{3}\right)$ in $G_{3}^{1}$ is defined by

$$
\langle A, B\rangle_{G_{3}^{1}}= \begin{cases}a_{1} b_{1}, & \text { if } \quad a_{1} \neq 0 \vee b_{1} \neq 0  \tag{2.3}\\ a_{2} b_{2}-a_{3} b_{3}, & \text { if } \quad a_{1}=0 \wedge b_{1}=0\end{cases}
$$

The pseudo-Galilean cross product is defined for $a=\left(a_{1}, a_{2}, a_{3}\right), b=\left(b_{1}, b_{2}, b_{3}\right)$ by

$$
a \wedge_{G_{3}^{1}} b=\left|\begin{array}{ccc}
0 & -e_{2} & e_{3} \\
a_{1} & a_{2} & a_{3} \\
b_{1} & b_{2} & b_{3}
\end{array}\right|
$$

A curve $\alpha(t)=(x(t), y(t), z(t))$ is admissible if it has no inflection points, no isotropic tangents or tangents or normals whose projections on the absolute plane would be light-like vectors. For an admissible curve $\alpha: I \subseteq R \rightarrow G_{3}^{1}$ the curvature $\kappa(t)$ and the torsion $\tau(t)$ are defined by

$$
\begin{equation*}
\kappa(t)=\frac{\sqrt{\left(y^{\prime \prime}(t)\right)^{2}-\left(z^{\prime \prime}(t)\right)^{2}}}{\left(x^{\prime}(t)\right)^{2}}, \quad \tau(t)=\frac{y^{\prime \prime}(t) z^{\prime \prime \prime}(t)-y^{\prime \prime \prime}(t) z^{\prime \prime}(t)}{\left|x^{\prime}(t)\right|^{5} \kappa^{2}(t)} \tag{2.4}
\end{equation*}
$$

expressed in components. Hence, for an admissible curve $\alpha: I \subseteq R \rightarrow G_{3}^{1}$ parameterized by the arc length $s$ with differential form $d s=d x$, given by

$$
\begin{equation*}
\alpha(t)=(x, y(s), z(s)) \tag{2.5}
\end{equation*}
$$

the formulas (2.5) have the following form

$$
\begin{equation*}
\kappa(s)=\sqrt{\left|\left(y^{\prime \prime}(s)\right)^{2}-\left(z^{\prime \prime}(s)\right)^{2}\right|}, \quad \tau(s)=\frac{y^{\prime \prime}(s) z^{\prime \prime \prime}(s)-y^{\prime \prime \prime}(s) z^{\prime \prime}(s)}{\kappa^{2}(s)} \tag{2.6}
\end{equation*}
$$

The associated trihedron is given by

$$
\begin{align*}
T & =\alpha^{\prime}(s)=\left(1, y^{\prime}(s), z^{\prime}(s)\right)  \tag{2.7}\\
N & =\frac{1}{\kappa(s)} \alpha^{\prime \prime}(s)=\frac{1}{\kappa(s)}\left(0, y^{\prime \prime}(s), z^{\prime \prime}(s)\right) \\
B & =\frac{1}{\kappa(s)}\left(0, \epsilon z^{\prime \prime}(s), \epsilon y^{\prime \prime}(s)\right)
\end{align*}
$$

where $\epsilon=\mp 1$, chosen by criterion $\operatorname{det}(T, N, B)=1$, that means

$$
\left|\left(y^{\prime \prime}(s)\right)^{2}-\left(z^{\prime \prime}(s)\right)^{2}\right|=\epsilon\left(\left(y^{\prime \prime}(s)\right)^{2}-\left(z^{\prime \prime}(s)\right)^{2}\right)
$$

We derive an important relation

$$
\alpha^{\prime \prime \prime}(s)=\kappa^{\prime}(s) N(s)+\kappa(s) \tau(s) B(s) .
$$

The curve $\alpha$ given by (2.5) is timelike (resp. spacelike) if $N(s)$ is a spacelike(resp. timelike) vector. The principal normal vector or simply normal is spacelike if $\epsilon=1$ and timelike if $\epsilon=-1$. For derivatives of the tangent (vector) $T$, the normal $N$ and the binormal $B$, respectively, the following Serret-Frenet formulas [8, 9] hold

$$
\begin{align*}
T^{\prime} & =\kappa N  \tag{2.8}\\
N^{\prime} & =\tau B \\
B^{\prime} & =\tau N
\end{align*}
$$

Definition 2.1. Let $G_{3}^{1}$ be the 3 -dimensional pseudo-Galilean space with the standard inner product $\langle,\rangle_{G_{3}^{1}}$. If there exists a corresponding relationship between the
admissible curves $\alpha$ and $\beta$ such that, at the corresponding points of the admissible curves, the principal normal lines of $\beta$ coincides with the binormal lines of $\alpha$, then $\beta$ is called an admissible Mannheim curve, and $\alpha$ a Mannheim partner curve of $\beta$. The pair $\{\alpha, \beta\}$ is said to be a Mannheim pair [8].

Definition 2.2. An admissible Mannheim curve $\beta\left(s^{\star}\right), s^{\star} \in I$ is a $C^{\infty}$ regular curve with non-zero curvature for which there exists another (different) $C^{\infty}$ regular curve $\alpha(s)$ where $\alpha(s)$ is of class $C^{\infty}$ and $\alpha^{\prime}(s) \neq 0$ ( $s$ being the arc length of $\alpha(s)$ only), also with non-zero curvature, in bijection with it in such a manner that the principal normal to $\beta\left(s^{\star}\right)$ and the binormal to $\alpha(s)$ at each pair of corresponding points coincide with the line joining the corresponding points. The curve $\alpha(s)$ is called a Mannheim conjugate of $\beta\left(s^{\star}\right)$.

Definition 2.3. An admissible Frenet-Mannheim curve $\beta\left(s^{\star}\right)$ (briefly called a $F M$ curve) is a $C^{\infty}$ Frenet curve for which there exists another $C^{\infty}$ Frenet curve $\alpha(s)$, where $\alpha(s)$ is of class $C^{\infty}$ and $\alpha^{\prime}(s) \neq 0$, in bijection with it so that, by suitable choice of the Frenet frames the principal normal vector $N_{\beta}\left(s^{\star}\right)$ and binormal vector $B_{\alpha}(s)$ at corresponding points on $\beta\left(s^{\star}\right), \alpha(s)$, both lie on the line joining the corresponding points. The curve $\alpha(s)$ is called a FM conjugate of $\beta\left(s^{\star}\right)$.

Definition 2.4. An admissible weakened Mannheim curve $\beta\left(s^{\star}\right), s^{\star} \in I^{\star}$ (briefly called a $W M$ curve) is a $C^{\infty}$ regular curve for which there exists another $C^{\infty}$ regular curve $\alpha(s), s \in I$, where $s$ is the arclength of $\alpha(s)$, and a homeomorphism $\sigma: I \rightarrow I^{\star}$ such that
(i) There exist two (disjoint) closed subsets $Z, N$ of $I$ with void interiors such that $\sigma \in C^{\infty}$ on $I \backslash N,\left(\frac{d s^{\star}}{d s}\right)=0$ on $Z, \sigma^{-1} \in C^{\infty}$ on $\sigma(I \backslash Z)$ and $\left(\frac{d s}{d s^{\star}}\right)=0$ on $\sigma(N)$;
(ii) The line joining corresponding points $s, s^{\star}$ of $\alpha(s)$ and $\beta\left(s^{\star}\right)$ is orthogonal to $\alpha(s)$ and $\beta\left(s^{\star}\right)$ at the points $s, s^{\star}$ respectively, and is along the principal normal to $\beta\left(s^{\star}\right)$ or $\alpha(s)$ at the points $s, s^{\star}$ whenever it is well defined.

The curve $\alpha(s)$ is called a $W M$ conjugate of $\beta\left(s^{\star}\right)$.
Thus for a WM curve we not only drop the requirement of $\alpha(s)$ being a Frenet curve, but also allow $\left(\frac{d s^{*}}{d s}\right)$ to be zero on a subset with void interior $\left(\frac{d s^{\star}}{d s}\right)=0$ on an interval would destroy the injectivity of the mapping $\sigma$. Since $\left(\frac{d s^{\star}}{d s}\right)=0$ implies that $\left(\frac{d s}{d s^{\star}}\right)$ does not exist, the apparently artificial requirements in (i) are in fact quite natural.

It is clear that an admissible Mannheim curve is necessarily a FM curve, and a FM curve is necessarily a WM curve. It will be proved in Theorem4.3 that under certain conditions a WM curve is also a FM curve.

## 3 Frenet-Mannheim Curves

In this section we study the structure and characterization of FM curves. We begin with a lemma, the method used in which is classical.

Lemma 3.1. Let $\beta\left(s^{\star}\right)$, $s^{\star} \in I^{\star}$ be a $F M$ curve and $\alpha(s)$ a $F M$ conjugate of $\beta\left(s^{\star}\right)$. Let

$$
\begin{equation*}
\beta\left(s^{\star}\right)=\alpha(s)+\lambda(s) B_{\alpha}(s) \tag{3.1}
\end{equation*}
$$

Then the distance $|\lambda|$ between corresponding points of $\alpha(s), \beta\left(s^{\star}\right)$ is constant, and there is a constant angle $\theta$ such that $\left\langle T_{\alpha}, T_{\beta}\right\rangle=\cos \theta$ and
(i) $\sinh \theta=\lambda \tau_{\alpha} \cosh \theta$;
(ii) $\sinh \theta=\lambda \tau_{\beta} \cosh \theta$;
(iii) $\cosh ^{2} \theta=1$;
(iv) $\sinh ^{2} \theta=\lambda^{2} \tau_{\alpha} \tau_{\beta}$.

Proof. From (3.1) it follows that

$$
\lambda(s)=\left\langle\beta\left(s^{\star}\right)-\alpha(s), B_{\alpha}(s)\right\rangle
$$

is of class $C^{\infty}$. Differentiation of (3.1) with respect to $s$ gives

$$
\begin{equation*}
T_{\beta} \frac{d s^{\star}}{d s}=T_{\alpha}+\lambda^{\prime} B_{\alpha}+\lambda \tau_{\alpha} N_{\alpha} \tag{3.2}
\end{equation*}
$$

Since by hypothesis we have $B_{\alpha}=\epsilon N_{\beta}$ with $\epsilon= \pm 1$, scalar multiplication of (3.2) by $B_{\alpha}$ gives

$$
\lambda^{\prime}=0,
$$

then we have $\lambda$ is a constant function. Therefore

$$
\begin{equation*}
T_{\beta} \frac{d s^{\star}}{d s}=T_{\alpha}+\lambda \tau_{\alpha} N_{\alpha} \tag{3.3}
\end{equation*}
$$

But by definition of FM curve we have $\frac{d s^{\star}}{d s} \neq 0$, so that $T_{\beta}$ is $C^{\infty}$ function of $s$. Hence

$$
\left\langle T_{\alpha}, T_{\beta}\right\rangle_{G_{3}^{1}}^{\prime}=\kappa_{\alpha}\left\langle N_{\alpha}, T_{\beta}\right\rangle_{G_{3}^{1}}+\frac{d s^{\star}}{d s} \kappa_{\beta}\left\langle T_{\alpha}, N_{\beta}\right\rangle_{G_{3}^{1}}=0 .
$$

Consequently $\left\langle T_{\alpha}, T_{\beta}\right\rangle$ is constant, and there exists a constant angle $\theta$ such that

$$
\begin{equation*}
T_{\beta}=T_{\alpha} \cosh \theta+N_{\alpha} \sinh \theta \tag{3.4}
\end{equation*}
$$

Taking the vector product of (3.3) and (3.4), we obtain

$$
\sin \theta=\lambda \tau_{\alpha} \cosh \theta
$$

which is $(i)$. Now write

$$
\alpha(s)=\beta\left(s^{\star}\right)-\epsilon \lambda(s) N_{\beta}(s) .
$$

Therefore

$$
\begin{equation*}
T_{\alpha}=\frac{d s^{\star}}{d s}\left[T_{\beta}-\lambda \epsilon \tau_{\beta} B_{\beta}\right] . \tag{3.5}
\end{equation*}
$$

On the other hand, equation (3.4) gives

$$
B_{\beta}=T_{\beta} \wedge_{G_{3}^{1}} N_{\beta}=\epsilon N_{\alpha} \cosh \theta
$$

Using (3.4) again, we get

$$
\begin{equation*}
T_{\alpha}=T_{\beta} \cosh \theta-\epsilon B_{\beta} \sinh \theta \tag{3.6}
\end{equation*}
$$

Taking the vector product of (3.5) and (3.6), we obtain

$$
\sinh \theta=\lambda \tau_{\beta} \cosh \theta
$$

which is (ii). On the other hand, comparison of (3.3) and (3.4) gives

$$
\begin{gather*}
\frac{d s^{\star}}{d s} \cosh \theta=1,  \tag{3.7}\\
\frac{d s^{\star}}{d s} \sinh \theta=\lambda \tau_{\alpha} . \tag{3.8}
\end{gather*}
$$

Similarly (3.5), (3.6) give

$$
\begin{gather*}
\frac{d s^{\star}}{d s}=\cosh \theta  \tag{3.9}\\
\frac{d s^{\star}}{d s}\left(\lambda \tau_{\beta}\right)=\sinh \theta \tag{3.10}
\end{gather*}
$$

The properties (iii) and (iv) then easily follow from (3.7) and (3.9), (3.6) and (3.8) and (3.10).

Theorem 3.2. Let $\beta\left(s^{\star}\right)$, $s^{\star} \in I^{\star}$ be a $C^{\infty}$ Frenet curve with $\tau_{\beta}$ nowhere zero and satisfying the equation for constants $\lambda$ with $\lambda \neq 0$. Then $\beta\left(s^{\star}\right)$ is a non-planar FM curve.

$$
\begin{equation*}
\sinh \theta=\lambda \tau_{\beta} \cosh \theta \tag{3.11}
\end{equation*}
$$

Proof. Define the curve $\beta\left(s^{\star}\right)$ with position vector

$$
\beta\left(s^{\star}\right)=\alpha(s)+\lambda(s) B_{\alpha}(s)
$$

Then, denoting differentiation with respect to $s$ by a dash, we have

$$
\beta^{\prime}\left(s^{\star}\right)=T_{\alpha}+\lambda \tau_{\alpha} N_{\alpha} .
$$

Since $\tau_{\alpha} \neq 0$, it follows that $\beta\left(s^{\star}\right)$ is a $C^{\infty}$ regular curve. Then

$$
T_{\beta} \frac{d s^{\star}}{d s}=T_{\alpha}+\lambda \tau_{\alpha} N_{\alpha}
$$

Hence

$$
\frac{d s^{\star}}{d s}=\sqrt{1+\lambda^{2} \tau_{\alpha}^{2}}
$$

And, using (3.11)

$$
T_{\beta}=T_{\alpha} \cosh \theta+N_{\alpha} \sinh \theta
$$

notice that from (3.11) we have $\sinh \theta \neq 0$. Therefore

$$
\frac{T_{\beta}}{d s^{\star}} \frac{d s^{\star}}{d s}=\kappa_{\alpha} N_{\alpha} \cosh \theta+\tau_{\alpha} B_{\alpha} \sinh \theta
$$

Now define $N_{\beta}=\epsilon B_{\alpha}$,

$$
\kappa_{\beta}=\frac{\epsilon}{\frac{d s^{\star}}{d s}} \tau_{\alpha} \sinh \theta
$$

These are $C^{\infty}$ functions of $s$ (and hence of $s^{\star}$ ), and

$$
\frac{T_{\beta}}{d s^{\star}}=\kappa_{\beta} N_{\beta}
$$

Further define $B_{\beta}=T_{\beta} \wedge_{G_{3}^{1}} B_{\alpha}$ and $\tau_{\beta}=\left\langle\frac{B_{\beta}}{d s^{\star}}, N_{\beta}\right\rangle_{G_{3}^{1}}$. These are also $C^{\infty}$ functions on $I^{\star}$. It is then easy to verify that with the frame $\left\{T_{\beta}, N_{\beta}, B_{\beta}\right\}$ and the functions $\kappa_{\beta}, \tau_{\beta}$, the curve $\beta\left(s^{\star}\right)$ becomes a $C^{\infty}$ Frenet curve. But $B_{\alpha}$ and $N_{\beta}$ lie on the line joining corresponding points of $\alpha(s)$ and $\beta\left(s^{\star}\right)$. Thus $\beta\left(s^{\star}\right)$ is a FM curve and $\alpha(s)$ a FM conjugate of $\beta\left(s^{\star}\right)$.

Lemma 3.3. A necessary and sufficient condition for a $C^{\infty}$ regular curve $\beta$ to be a FM curve with a FM conjugate. Then $\beta$ should be either a line or a non-planar circular helix.

Proof. $(\Rightarrow)$ : Let $\beta$ have a FM conjugate $\alpha$ which is a line. Then $\kappa_{\alpha}=0$. Using Lemma 3.1, (iii) and (i), (ii), we have

$$
\begin{equation*}
\cosh ^{2} \theta=1 \tag{3.12}
\end{equation*}
$$

and then

$$
\begin{gather*}
\cosh ^{2} \theta \sin \theta=\lambda \tau_{\beta} \cosh \theta  \tag{3.13}\\
\sinh \theta=\lambda \tau_{\alpha} \cosh \theta \tag{3.14}
\end{gather*}
$$

From (3.14) it follows that $\cosh \theta \neq 0$. Hence (3.13) is equivalent to

$$
\begin{equation*}
\lambda \tau_{\beta}=\cosh \theta \sinh \theta \tag{3.15}
\end{equation*}
$$

Case 1. $\sinh \theta=0$. Then $\cosh \theta= \pm 1$, so that (3.12) implies that $\kappa_{\beta}=0$, and $\beta$ is a line. We note also that (3.15) implies that $\tau_{\beta}=0$.

Case 2. $\sinh \theta \neq 0$. Then $\cosh \theta \neq \pm 1$, and (3.12), (3.15) imply that $\kappa_{\beta}, \tau_{\beta}$ are non-zero constants, and $\beta$ is a non-planar circular helix.
$(\Leftarrow)$ : If $\beta$ is a non-planar circular helix

$$
\beta=(a s, b \cosh s, b \sinh s),
$$

we may take

$$
N_{\beta}=(0, \cosh s, \sinh s) .
$$

Now put $\lambda=b$, then the curve $\beta$ with

$$
\beta=\alpha+\lambda B_{\alpha}
$$

will be a line along the $x$-axis, and can be made into a FM conjugate of $\beta$ if $N_{\beta}$ is defined to be equal to $B_{\alpha}$.

Theorem 3.4. Let $\beta\left(s^{\star}\right)$ be a plane $C^{\infty}$ Frenet curve with zero torsion and whose curvature is either bounded below or bounded above. Then $\beta$ is a FM curve, and has FB conjugates which are plane curves.

Proof. Let $\beta$ be a curve satisfying the conditions of the hypothesis. Then there are non-zero numbers $\lambda$ such that $\kappa_{\beta}<-\frac{1}{\lambda}$ on $I$ or $\kappa_{\beta}>-\frac{1}{\lambda}$ on $I$. For any such $\lambda$, consider the plane curve $\alpha$ with position vector

$$
\alpha=\beta-\lambda N_{\beta} .
$$

Then

$$
T_{\alpha}=T_{\beta}
$$

It is then a straightforward matter to verify that $\alpha$ is a FM conjugate of $\beta$.

## 4 Weakened Mannheim Curves

Definition 4.1. Let $D$ be a subset of a topological space $X$. A function on $X$ into a set $Y$ is said to be $D$-piecewise constant if it is constant on each component of $D$.

Lemma 4.2. Let $X$ be a proper interval on the real line and $D$ an open subset of $X$. Then a necessary and sufficient condition for every continuous, $D$-piecewise constant real function on $X$ to be constant is that $X \backslash D$ should have empty dense-in-itself kernel.

We notice that if $D$ is dense in $X$, any $C^{1}$ and $D$-piecewise constant real function on $X$ must be constant, even if $D$ has non-empty dense-in-itself kernel.

Theorem 4.3. $A$ WM curve for which $N$ and $Z$ have empty dense-in-itself kernels is a FM curve.

Proof. Let $\beta\left(s^{\star}\right), s^{\star} \in I^{\star}$ be a WM curve and $\alpha(s), s \in I$ a WM conjugate of $\beta\left(s^{\star}\right)$. It follows from the definition that $\alpha(s)$ and $\beta\left(s^{\star}\right)$ each has a $C^{\infty}$ family of tangent vectors $T_{\beta}\left(s^{\star}\right), T_{\alpha}(s)$. Let

$$
\begin{equation*}
\beta(s)=\beta(\sigma(s))=\alpha(s)+\lambda(s) B_{\alpha}(s) \tag{4.1}
\end{equation*}
$$

where $B_{\alpha}(s)$ is some unit vector function and $\lambda(s) \geq 0$ is some scalar function. Let $D=I \backslash N, D^{\star}=I^{\star} \backslash \sigma(Z)$. Then $s^{\star}(s) \in C^{\infty}$ on $D^{\star}$.

Step 1. To prove $\lambda=$ constant.
Since $\lambda=\|\beta(s)-\alpha(s)\|$, it is continuous on $I$ and is of class $C^{\infty}$ on every interval of $D$ on which it is nowhere zero. Let $P=\{s \in I: \lambda(s) \neq 0\}$ and $X$ any component of $P$. Then $P$, and hence also $X$, is open in $I$. Let $L$ be any component interval of $X \cap D$. Then on $L, \lambda(s)$ and $B_{\alpha}(s)$ are of class $C^{\infty}$, and from (4.1) we have

$$
\beta^{\prime}(s)=\alpha^{\prime}(s)+\lambda^{\prime}(s) B_{\alpha}(s)+\lambda(s) B_{\alpha}^{\prime}(s) .
$$

Now by definition of a WM curve we have $\left\langle\alpha^{\prime}(s), B_{\alpha}(s)\right\rangle_{G_{3}^{1}}=0=\left\langle\beta^{\prime}\left(s^{\star}\right), B_{\alpha}(s)\right\rangle_{G_{3}^{1}}$. Hence, using the identity $\left\langle B_{\alpha}^{\prime}(s), B_{\alpha}(s)\right\rangle_{G_{3}^{1}}=0$, we have

$$
0=\lambda^{\prime}(s)\left\langle B_{\alpha}(s), B_{\alpha}(s)\right\rangle_{G_{3}^{1}}
$$

Therefore $\lambda=$ constant on $L$.
Hence $\lambda$ is constant on each interval of the set $X \cap D$. But by hypothesis $X \backslash D$ has empty dense-in-itself kernel. It follows from Lemma 3.3 that $\lambda$ is constant (and non-zero) on $X$. Since $\lambda$ is continuous on $I, X$ must be closed in $I$. But $X$ is also open in $I$. Therefore by connectedness we must have $X=I$, that is, $\lambda$ is constant on $I$.

Step 2. To prove the existence of two frames

$$
\left\{T_{\alpha}(s), N_{\alpha}(s), B_{\alpha}(s)\right\},\left\{T_{\beta}\left(s^{\star}\right), N_{\beta}\left(s^{\star}\right), B_{\beta}\left(s^{\star}\right)\right\}
$$

which are Frenet frames for $\alpha(s), \beta\left(s^{\star}\right)$ on $D, D^{\star}$ respectively.
Since $\lambda$ is a non-zero constant, it follows from (4.1) that $B_{\alpha}(s)$ is continuous on $I$ and $C^{\infty}$ on $D$, and is always orthogonal to $T_{\alpha}(s)$. Now define $B_{\alpha}(s)=$ $T_{\alpha}(s) \wedge_{G_{3}^{1}} N_{\alpha}(s)$. Then $\left\{T_{\alpha}(s), N_{\alpha}(s), B_{\alpha}(s)\right\}$ forms a right-handed orthonormal frame for $\alpha(s)$ which is continuous on $I$ and $C^{\infty}$ on $D$.

Now from the definition of WM curve we see that there exists a scalar function $\kappa_{\beta}\left(s^{\star}\right)$ such that $T_{\beta}^{\prime}\left(s^{\star}\right)=\kappa_{\beta}\left(s^{\star}\right) N_{\beta}\left(s^{\star}\right)$ on $I^{\star}$. Hence $\kappa_{\beta}\left(s^{\star}\right)=\left\langle T_{\beta}^{\prime}\left(s^{\star}\right), N_{\beta}\left(s^{\star}\right)\right\rangle_{G_{3}^{1}}$ is continuous on $I^{\star}$ and $C^{\infty}$ on $D^{\star}$. Thus the first Frenet formula holds on $D^{\star}$. It is then straightforward to show that there exists a $C^{\infty}$ function $\tau_{\alpha}(s)$ on $D$ such that the Frenet formulas hold. Thus $\left\{T_{\alpha}(s), N_{\alpha}(s), B_{\alpha}(s)\right\}$ is a Frenet frame for $\alpha(s)$ on $D$.

Similarly there exists a right-handed orthonormal frame $\left\{T_{\beta}\left(s^{\star}\right), N_{\beta}\left(s^{\star}\right), B_{\beta}\left(s^{\star}\right)\right\}$ for $\beta\left(s^{\star}\right)$ which is continuous on $I^{\star}$ and is a Frenet frame for $\beta\left(s^{\star}\right)$ on $D^{\star}$. Moreover, we can choose

$$
B_{\alpha}(s)=N_{\beta}(\sigma(s))
$$

Step 3. To prove that $N=\emptyset, Z=\emptyset$. We first notice that on $D$ we have

$$
\left\langle T_{\beta}, T_{\alpha}\right\rangle_{G_{3}^{1}}^{\prime}=\left\langle\kappa_{\beta} N_{\beta} \frac{d s^{\star}}{d s}, T_{\alpha}\right\rangle_{G_{3}^{1}}+\left\langle T_{\beta}, \kappa_{\alpha} N_{\alpha}\right\rangle_{G_{3}^{1}}=0
$$

so that $\left\langle T_{\beta}, T_{\alpha}\right\rangle$ is constant on each component of $D$ and hence on $I$ by Lemma 4.2. Consequently there exists a angle $\theta$ such that

$$
T_{\beta}=T_{\alpha} \cosh \theta+N_{\alpha} \sinh \theta
$$

Further,

$$
B_{\alpha}(s)=N_{\beta}(\sigma(s))
$$

and so

$$
B_{\beta}\left(s^{\star}\right)=-T_{\alpha} \sinh \theta+N_{\alpha} \cosh \theta
$$

Thus $\left\{T_{\beta}\left(s^{\star}\right), N_{\beta}\left(s^{\star}\right), B_{\alpha}(s)\right\}$ are also of class $C^{\infty}$ on $D$. On the other hand $\left\{T_{\beta}\left(s^{\star}\right), N_{\beta}\left(s^{\star}\right), B_{\beta}\left(s^{\star}\right)\right\}$ are of class $C^{\infty}$ with respect to $s^{\star}$ on $D^{\star}$. Writing (4.1) in the form

$$
\alpha=\beta-\lambda N_{\beta}
$$

and differentiating with respect to $s$ on $D \cap \sigma^{-1}\left(D^{\star}\right)$, we have

$$
T_{\alpha}=\frac{d s^{\star}}{d s}\left[T_{\beta}+\lambda \tau_{\beta} B_{\beta}\right]
$$

But

$$
T_{\alpha}=T_{\beta} \cosh \theta-B_{\beta} \sinh \theta
$$

Hence

$$
\begin{equation*}
\frac{d s^{\star}}{d s}=\cosh \theta \text { and } \lambda \tau_{\beta}=\sinh \theta \tag{4.2}
\end{equation*}
$$

Since $\kappa_{\beta}\left(s^{\star}\right)=\left\langle T_{\beta}^{\prime}, N_{\beta}\right\rangle_{G_{3}^{1}}$ is defined and continuous on $I^{\star}$ and $\sigma^{-1}\left(D^{\star}\right)$ is dense, it follows by continuity that (4.2) holds throughout $D$. If $\cosh \theta \neq 0$ then (4.2) implies that $\frac{d s^{\star}}{d s} \neq 0$ on $D$. Hence $Z=\emptyset$. Similarly $N=\emptyset$.

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